

# Observational Astronomy - Lecture 10

## Galaxies - Structure, Types, and Evolution

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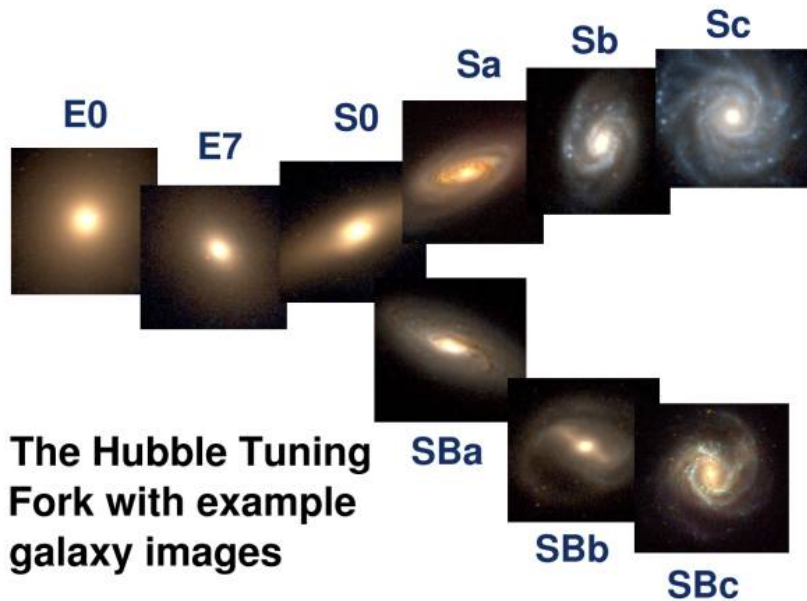
April 21, 2014

# Galaxy Basics - 1

- 1 Galaxies are large groups of stars held together by gravity.
- 2 They vary widely in size:
  - Large galaxies like the Milky way have  $10^{11} - 10^{12}$  stars.
  - Small dwarf galaxies may have only a few million stars.
- 3 They come in a wide variety of shapes, but there are two main types:
  - Spiral galaxies, which are blue in color because they are still forming stars.
  - Elliptical galaxies, which are no longer forming many stars and are therefore red in color.
  - The star formation rate of the universe peaked long ago and is declining.

- ④ We believe that all galaxies have a Super-Massive Black Hole (SMBH) at their centers.
- ⑤ Galaxies grow through collisions and mergers.
- ⑥ Quasars (Quasi-stellar objects) are galaxies whose SMBH is accreting gas at a rapid rate:
  - Because the hot gas emits radiation, they are extremely luminous objects.
  - They were much more numerous in the early universe
  - There are no nearby quasars.

# The Hubble Tuning Fork Diagram



# NGC 1300: A Typical Barred Spiral

Barred Spiral Galaxy NGC 1300



Hubble  
Heritage

# Red Ellipticals vs Blue Spirals



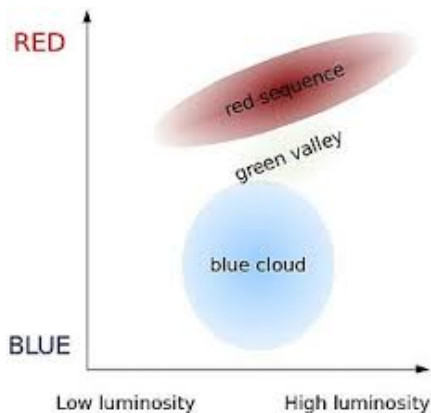
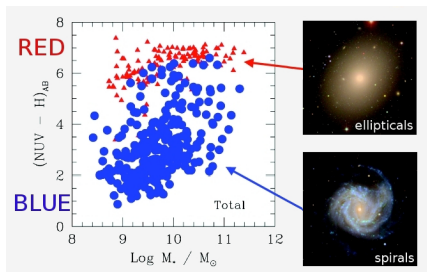
An Elliptical Galaxy.  
"Red and Dead"



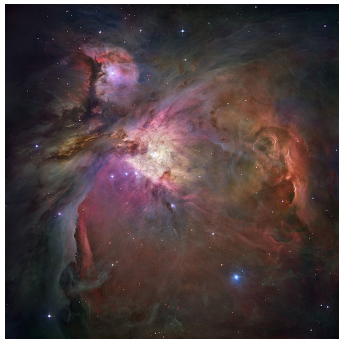
A spiral galaxy.  
Still forming hot blue stars.

- Spiral galaxies like the Milky Way are still forming new stars.
- Elliptical galaxies have used up their gas and are no longer forming large numbers of stars.

# The Red Sequence, the Blue Cloud, and the Green Valley



# H-II Regions

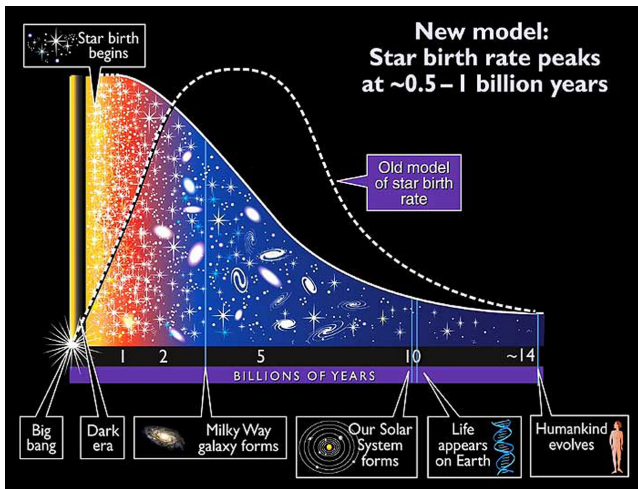


The Orion Nebula is a nearby H-II region.

- The gas in H-II regions is heated to luminescence by embedded hot young stars.
- H-II regions are indicative of star formation.
- The pink color is due to emission from ionized hydrogen ( $H - \alpha$ ).



# The Star Formation History of the Universe



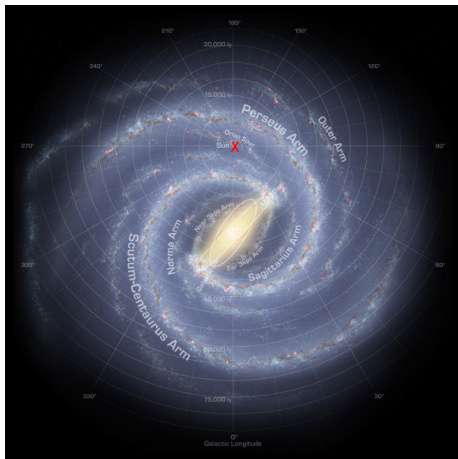
While the details are uncertain, it is clear that the star formation rate of the universe has dropped significantly

# The Milky Way as seen from Mt. Haleakala



The Milky Way is visible with the naked eye from a dark-sky site.

# A Map of the Milky Way



This is what we believe Milky Way would look like if seen from outside.

# NGC 6744: A Milky Way Look-alike



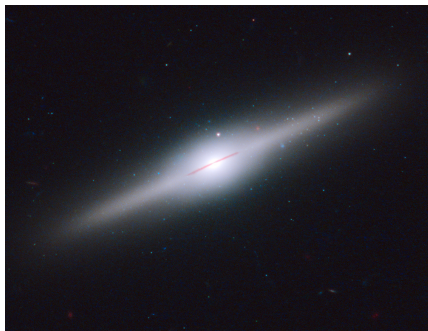
# Gas, Dust, and Dark Matter - the Composition of Galaxies

- As we will study in the coming weeks, we believe  $\approx 80 - 90\%$  of the mass of galaxies is in the form of “Dark Matter”. This matter is not made up of ordinary matter (i.e. it is not made up of atoms).
- The ordinary matter (which astronomers typically call “baryonic” matter), has the following components:
  - Stars and planets - these account for perhaps 10-50% of the ordinary matter.
  - Gas in the interstellar medium - this is most of the ordinary matter.
  - “Dust” - these are small solid particles made up of heavier elements.

# Dust in Galaxies

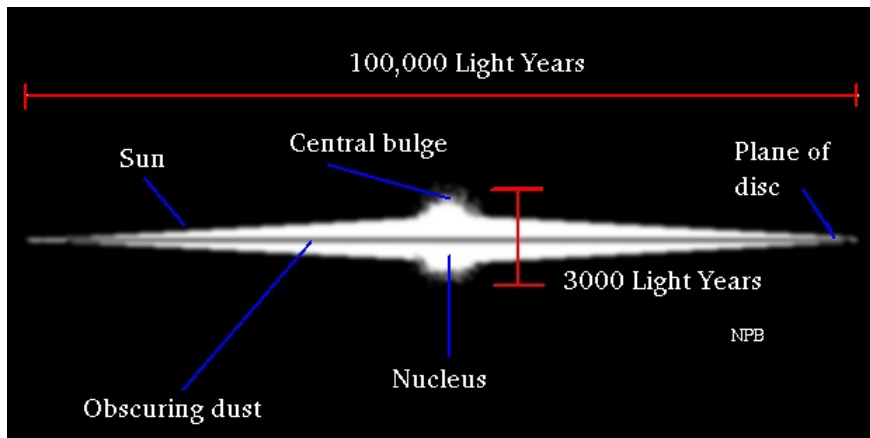
- Dust particles scatter light and block our observations.
- The scattering rate of dust particles is proportional to  $\frac{1}{\text{Wavelength}^4}$ .
- Using long wavelengths (infrared and radio), we can see through the dust clouds.
- The dust in spiral galaxies is concentrated in the galactic plane.
- We cannot see the center of our galaxy in visible light - we need to use infrared and radio telescopes.

## Edge-on Spiral Galaxies



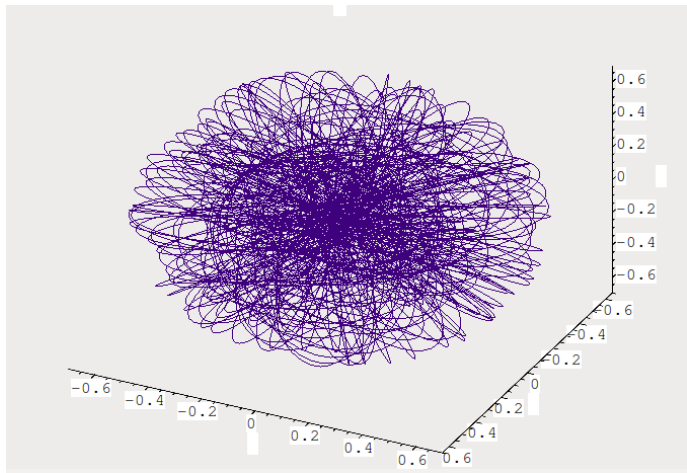
In Edge-on galaxies, we can clearly see the dust lane in the plane of the galaxy.

# A Milky Way Edge-On Schematic



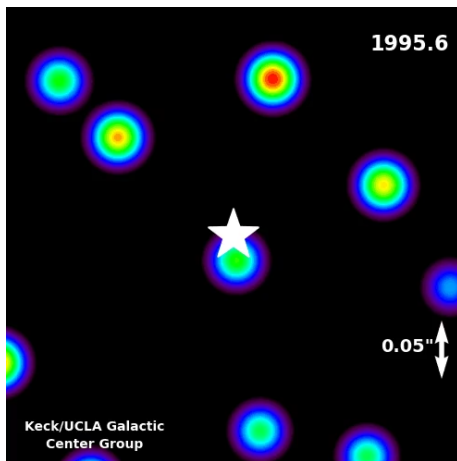


# A Typical Stellar Orbit in a Galaxy



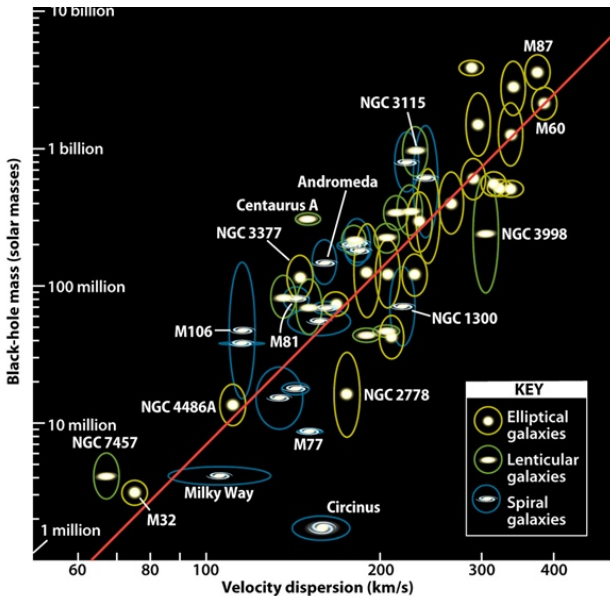
Unlike orbits in the Solar System, galactic orbits are typically not closed.

# Movie of Stars Orbiting our Central Black Hole



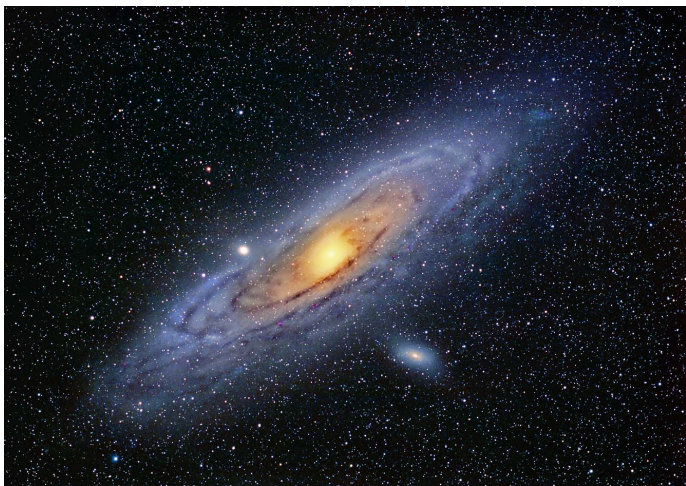
This movie shows stars orbiting our central black hole. From Kepler's laws, we can calculate that the black hole has a mass of about 4 million times the Sun.

# Galaxy Mass is correlated with the mass of the SMBH



ASTRONOMY: ROEN KELLY, AFTER KAYHAN GÜLTEKIN, ET AL.

# The Andromeda Galaxy (M31) - The Nearest Large Galaxy



This galaxy is visible with the naked eye from a dark-sky site  
The Milky Way and Andromeda will merge in  $\approx 4$  billion years.

# The Large and Small Magellanic Clouds - Nearby Galaxies



*Akira Fujii/David Malin Images*

# A “Warped” Galaxy

Galaxy ESO 510-G13



Hubble  
Heritage

# Interacting Galaxies - 1



# Interacting Galaxies - 2



Spiral Galaxy Pair NGC 3314

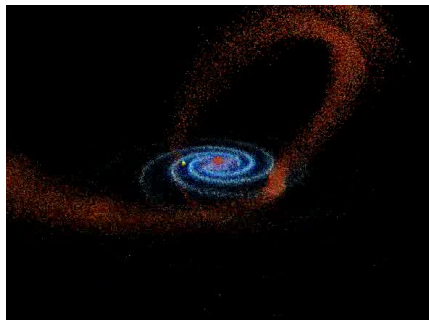
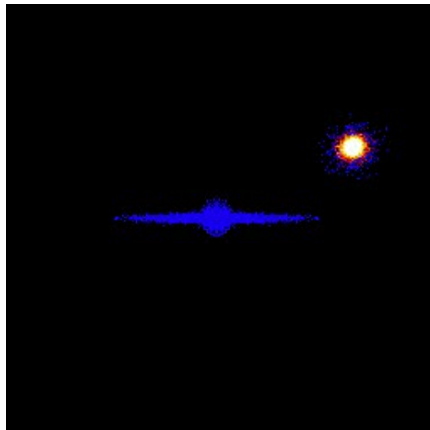


Hubble  
Heritage

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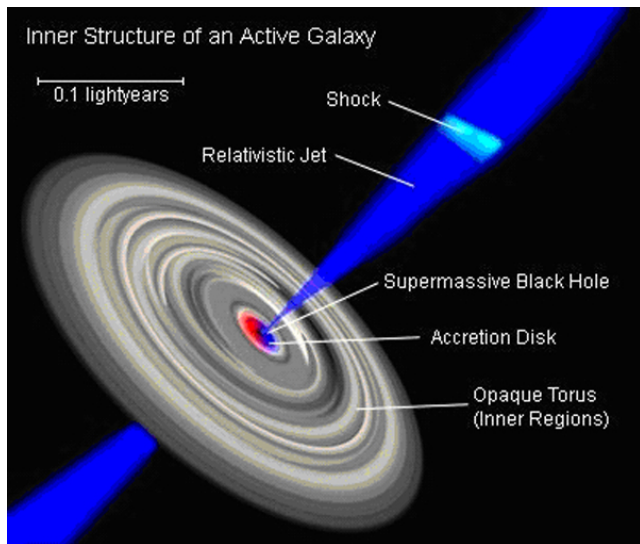
# Movies of Milky Way merging with the Sagittarius Dwarf Galaxy



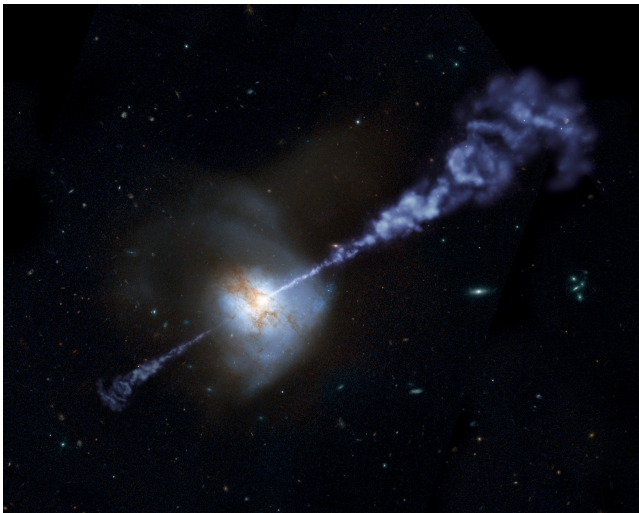
# Active Galaxies and Quasars

- Quasars (QUASi-stellAr-Radio Sources) were initially discovered as very distant objects which appeared star-like.
- Over time, we understood that these are large galaxies whose SMBH is accreting large quantities of gas and shining very brightly across a wide range of wavelengths.
  - They can be 100's to 1000's of times more luminous than the Milky Way.
- They come in a wide variety of sizes, luminosities, and appearances, depending on:
  - The size of the SMBH.
  - How rapidly the SMBH is accreting gas.
  - The angle at which we are viewing it.
- Quasars, Blazars, Active Galactic Nuclei (AGN's) are all names used to describe these objects.

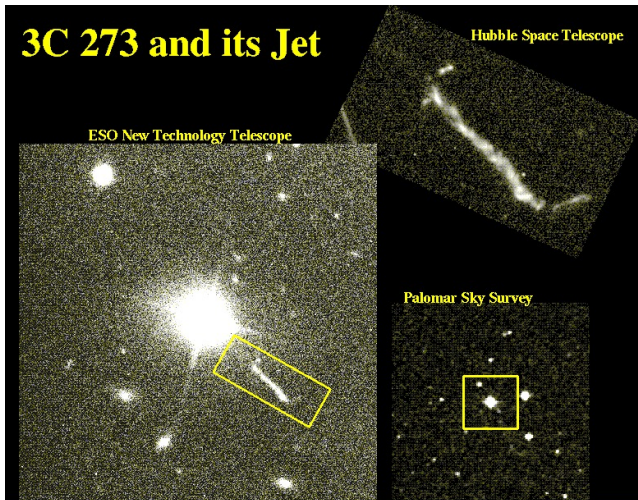
# Schematic of an Active Galactic Nucleus



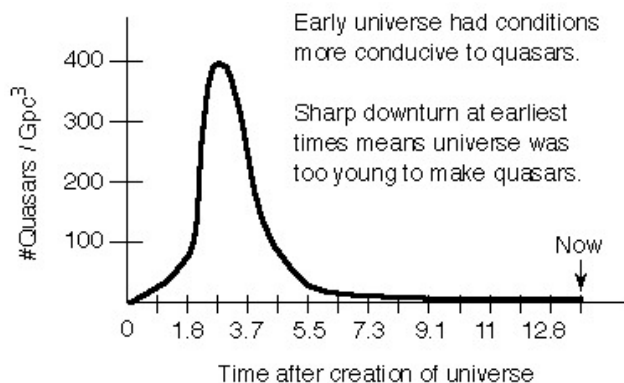
# The Nearest Active Galaxy - M87 in the Virgo Cluster



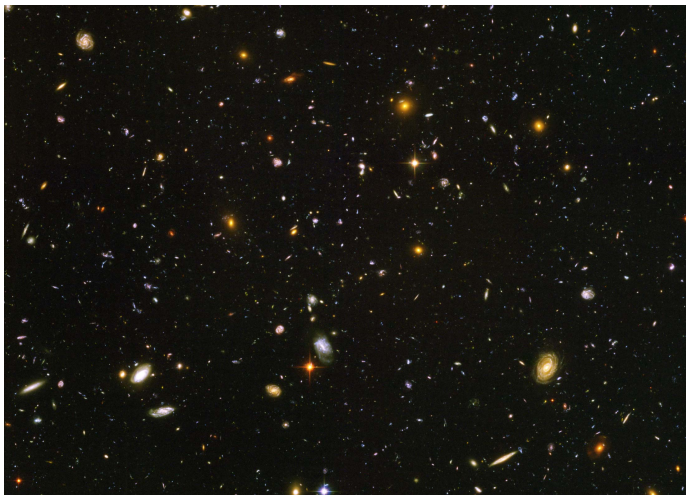
# The Quasar 3C273



## Quasars were much more common in the early Universe



# The Hubble Ultra Deep Field



- This is a tiny patch of sky, about 1/10 the size of the moon.
- The observable universe contains at least  $10^{11}$  galaxies!

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